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# $N$-particle dynamics of the Euler equations for planar diffeomorphisms 

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#### Abstract

The Euler equations associated with diffeomorphism groups have received much recent study because of their links with fluid dynamics, computer vision, and mechanics. In this article, we consider the dynamics of $N$ point particles or 'blobs' moving under the action of the Euler equations associated with the group of diffeomorphisms of the plane in a variety of different metrics. This dynamical system is already in widespread use in the field of image registration, where the point particles correspond to image landmarks, but its dynamical behavior has not previously been studied. The 2-body problem is always integrable, and we analyze its phase portrait under different metrics. In particular, we show that 2-body capturing orbits (in which the distances between the particles tend to 0 as $t \rightarrow \infty$ ) can occur when the kernel is sufficiently smooth and the relative initial velocity of the particles is sufficiently large. We compute the dynamics of these 'dipoles' with respect to other test particles, and supplement the calculations with simulations for larger $N$ that illustrate the different regimes.


## 1. Introduction

The Euler equations for planar diffeomorphisms are

$$
\begin{equation*}
\dot{m}+u \cdot \nabla m+\nabla u^{T} \cdot m+m(\operatorname{div} u)=0, \tag{1}
\end{equation*}
$$

where $\dot{m}$ denotes differentiation with respect to time, $u(x, t)\left(u, x \in \mathbb{R}^{n}, t \in \mathbb{R}\right)$ is a velocity field, and $m(x, t)$ its associated momentum. The velocity $u$ and momentum $m$ are related by

$$
\begin{equation*}
m=\mathcal{A} u \tag{2}
\end{equation*}
$$

where $\mathcal{A}$ is an elliptic operator (e.g., $\left.\mathcal{A}=\left(1-\nabla^{2}\right)^{k}\right)$ called the inertia operator.

[^0]Euler equations such as equations (1) and (2) have a natural geometric origin, which we sketch here although these details are not needed in the article. Let $\mathfrak{G}$ be a Lie or diffeomorphism group with Lie algebra $\mathfrak{g}$. Let $\mathfrak{G}$ be equipped with a left- or right-invariant metric that restricts to a metric $\langle\langle\rangle$,$\rangle on \mathfrak{g}$. Typically, this metric is defined by a linear inertia operator $\mathcal{A}: \mathfrak{g} \rightarrow \mathfrak{g}^{*}$ via

$$
\begin{equation*}
\langle\langle u, v\rangle\rangle:=\langle u, \mathcal{A} v\rangle . \tag{3}
\end{equation*}
$$

The geodesic equation on $T \mathfrak{G}$ can be reduced to give a noncanonical Hamiltonian evolution equation on $\mathfrak{g}$ called the Euler-Poincaré equation, or transferred (via the Legendre transform, $\mathfrak{g} \rightarrow \mathfrak{g}^{*}$, which in this case is $u \rightarrow m:=\mathcal{A} u$ ) to a Lie-Poisson system on $\mathfrak{g}^{*}$ called the Euler equations

$$
\begin{equation*}
\dot{m}= \pm \mathrm{ad}_{\mathcal{A}^{-1} m}^{*} m, \tag{4}
\end{equation*}
$$

where the sign is + for left- and - for right-invariant metrics.
The most famous Euler equation on a Lie group is the equation of motion of a free rigid body $(\mathfrak{G}=S O(3), m=$ body angular momentum). An infinite-dimensional example is the Landau-Lifshitz equation on the loop group $\mathfrak{G}=C^{\infty}(\mathbb{R}, S O(3))$. Examples on diffeomorphism groups include the Euler fluid equations on the group $\mathfrak{G}=\operatorname{Diff}_{\text {vol }}\left(\mathbb{R}^{n}\right)$ of volume-preserving diffeomorphisms with respect to the $L^{2}$ metric ( $m=$ vorticity) [5], the Camassa-Holm equation $\left(\mathfrak{G}=\operatorname{Diff}\left(\mathbb{S}^{1}\right), H^{1}\right.$ metric) [6], and the second-grade fluid equations $\left(\mathfrak{G}=\operatorname{Diff}_{\text {vol }}\left(\mathbb{R}^{n}\right)\right.$, $H^{1}$ metric) [16].

Equations (1) and (2) are the Euler equations for $\mathfrak{G}=\operatorname{Diff}\left(\mathbb{R}^{n}\right)$ with right-invariant metric (see [7, 15] for further details). It has been used recently (as the averaged template matching equations $[13,15]$ and the geodesic interpolating clamped-plate spline [12]) to align groups of medical images for use in automatic diagnosis. Pairs of images are smoothly interpolated by diffeomorphisms satisfying equation (1). Various metrics have been used, including Gaussian [8] and $H^{k}$ [12] metrics.

A striking feature of Euler equations on diffeomorphism groups is that they admit (formally, at least) exact solutions in which the momentum is concentrated at a finite set of points, which we call particles. For fluid equations these are point vortices, which are widely studied both in their own right and as a means of approximating the evolution of smooth or other vorticities. For the full diffeomorphism group these point particle solutions are used in the image registration problem $[18,12]$ and in this article we undertake an initial study of their dynamics. At first sight it is remarkable that a PDE should have such a finite-dimensional reduction, unrelated to symmetry or integrability. (Such a reduction, however, is not unique to Euler equations. On the other hand, there are many particle-based numerical methods, collectively known as Smoothed Particle Hydrodynamics [14], which are not exact solutions of the PDE they approximate.)

For the 2D and 3D Euler fluid equations, convergence of the point vortex solutions to solutions for smooth initial data has been established [1]. The speed of convergence can be improved by smoothing out the point vortices to vortex blobs [4], even though the (e.g., Gaussian) blobs are no longer an exact solution of the Euler equations. Instead, their evolution can be regarded as that of delta-functions under a slightly different inertia operator. Altering the shape of the blob is equivalent
to altering the metric; see section 3. In this way we are lead to consider a wide class of metrics. For an introduction to many-body problems, particularly arising in Newtonian and relativistic mechanics, see [2]. Point vortices defined over Dirac delta-functions, and the regularised version, vortex blob methods, are introduced in [4]; see [10] for a review of their use.

In contrast to the Euler fluid case, point particle solutions (as we term the solutions of equations (1) and (2) for diffeomorphisms of the plane) do not appear to evolve naturally from smooth initial data, and they may not constitute stable initial data. (Instead, bump-like initial data appears to evolve towards momentum sheets, not particles [8, 9].) Nevertheless, their dynamics are interesting and worthy of study not just because of their application in image registration, but also because they are a mathematically natural class of $N$-body systems that has never before been studied and they can be used as a numerical method to approximate smooth initial data, just as point vortices can. Indeed, convergence of this method has recently been established in the 1D, $H^{1}$ metric case (the Camassa-Holm equation) [3].

The inverse of the inertia operator $\mathcal{A}$ is given by convolution with the Green's function $\mathbf{G}$ of $\mathcal{A}$, i.e., $u=\mathbf{G} * m$, where $*$ denotes convolution and $\mathcal{A} \mathbf{G}(x)=\delta(x) I$ for $x \in \mathbb{R}^{n}$. We shall only consider rotationally invariant and diagonal $\mathcal{A}$; in this case $\mathbf{G}(x)=G(\|x\|) I$ for a scalar function $G$ which we call the kernel of the metric.

We begin the article by explicitly computing the Hamiltonian for the case of $N$ particles, and computing the equations of motion in terms of the kernel. Following this, in section 3 we use delta-function particles and Helmholtz-style inertia operators to find the kernels $G_{k}$ corresponding to $H^{k}$ metrics. The limit of these, $H^{\infty}$, corresponds to a Gaussian kernel for delta-function particles, where the inertia operator tends to $\exp \left(-\varepsilon^{2} \nabla^{2}\right)$. We show how even when blobs are required, rather than particles, we can still consider delta-function particles through a change in the inertia operator.

We then shift our attention to the case $N=2$, and examine the simplest possible interaction of particles. This two body problem is always integrable. We derive the reduced Hamiltonian and consider how the particle dynamics change for different metrics. The most striking feature of the dynamics is the phenomenon of particle capture, in which particles cohere or stick together. In a typical capture orbit, the distance between two particles is $\mathcal{O}\left(e^{-a t}\right)$ as $t \rightarrow \infty$, but there is no actual collision. We identify the regimes of capture and (more familiar) scattering orbits. For different metrics, we also show the phase portraits of the dynamics, and the appearance of typical scattering orbits. The analysis shows that the phase portrait depends sensitively on the behavior of $G$ close to $r=0$. Only when the kernel $G$ is smooth at $r=0$ and the relative initial velocities of the particles are sufficiently large will the particles capture each other, otherwise they scatter. For the cases $k \leq 2$ the Green's function is not sufficiently smooth to typically allow capture orbits.

Following this, in section 5 we consider the case $k>2$, so that the kernel is sufficiently smooth to allow capture orbits for two particles, and show how the dynamics of a set of $N$ particles varies as the particles approach each other. We study the behavior of 'dipoles', as we term pairs of particles that reach the limiting state of capture, and show a typical orbit for a group of 20 particles. We conclude with a summary of the differences between the particle dynamics studied here and that of point vortices, and identify a number of questions that remain open.

## 2. $N$-particle systems

The Hamiltonian for the Euler equation (4) is the kinetic energy

$$
\begin{equation*}
H=\frac{1}{2} \int\langle\langle u, m\rangle\rangle \mathrm{d} x \mathrm{~d} y=\frac{1}{2} \int m \mathcal{A}^{-1} m \mathrm{~d} x \mathrm{~d} y . \tag{5}
\end{equation*}
$$

Under the particle ansatz

$$
\begin{equation*}
m(x, t)=\sum_{i=1}^{N} p_{i}(t) \delta\left(x-q_{i}(t)\right) \tag{6}
\end{equation*}
$$

$\left(p_{i}(t), q_{i}(t) \in \mathbb{R}^{2}\right.$ for each $i$ ) the Hamiltonian becomes

$$
\begin{equation*}
H=\frac{1}{2} \sum_{i, j=1}^{N} p_{i} \cdot p_{j} G\left(\left\|q_{i}-q_{j}\right\|\right) \tag{7}
\end{equation*}
$$

Solutions to (4) of the form (6) obey Hamilton's equations for (7), in which the components of $q_{i}$ and $p_{i}$ are canonically conjugate variables (see [11] for further details). Here $q_{1}, \ldots, q_{N}$ represent the positions of the $N$ particles, and $p_{1}, \ldots, p_{N}$ represent their momenta.

There are two immediate differences from the analogous system for point vortices in the Euler fluid equations: the number of degrees of freedom is $2 N$ instead of $N$, and, to get well-defined ODEs, the inertia operator $\mathcal{A}$ must be chosen so that $G(0)$ is finite. (For point vortices, $G(r)=-\log (r) /(2 \pi)$, but the infinite self-energy of each particle can be ignored.)

The equations of motion are

$$
\begin{align*}
& \dot{q}_{i}=\sum_{j=1}^{N} G\left(\left\|q_{i}-q_{j}\right\|\right) p_{j}, \\
& \dot{p}_{i}=-\sum_{\substack{j=1 \\
j \neq i}}^{N}\left(p_{i} \cdot p_{j}\right) G^{\prime}\left(\left\|q_{i}-q_{j}\right\|\right) \frac{q_{i}-q_{j}}{\left\|q_{i}-q_{j}\right\|} . \tag{8}
\end{align*}
$$

They have four conserved quantities: the energy $H$, the linear momentum $\sum_{i=1}^{N} p_{i}$, and the angular momentum $\sum_{i=1}^{n} q_{i} \times p_{i}$. These are sufficient to ensure the integrability of the two particle problem, which is discussed in section 4. However, the dynamics of these 2-body problems are very different from more familiar ones on flat configuration spaces, like the Kepler problem. Also note that equations (8) are not Galilean-invariant; their dynamics depends on the linear momentum.

## 3. Blobs and kernels

The Euler equation for $\operatorname{Diff}\left(\mathbb{R}^{n}\right)$ and inertia operator $\mathcal{A}$ is given by (1) and (2). Applications to fluid mechanics often consider a related equation, the regularized Euler equation, for both numerical and analytic purposes. The regularization can be viewed in several different ways, but all give the same equations of motion: (i) the delta-function particles of momentum (or vorticity) can be replaced by smooth blobs; (ii) the kernel can be smoothed; or (iii) the inertia operator can be modified.

First we derive an expression for the kernel for standard delta-function particles in $\mathbb{R}^{2}$. The Green's function $G(x)$ is defined by $\mathcal{A} G(x)=\delta(x)$. Taking Fourier transforms gives $\tilde{a}(\omega) \widetilde{G}(\omega)=1$, where $\tilde{a}$ is the Fourier symbol of the operator $\mathcal{A}$. Our main example will be the operators

$$
\begin{equation*}
\mathcal{A}_{k}:=\left(1-\alpha^{2} \nabla^{2}\right)^{k}, \tag{9}
\end{equation*}
$$

with Fourier symbol

$$
\begin{equation*}
\tilde{a}(\omega)=\left(1+\alpha^{2}\|\omega\|^{2}\right)^{k} \tag{10}
\end{equation*}
$$

For rotationally-invariant $\mathcal{A}$, we have that $\tilde{a}(\omega)$ is a function of $\|\omega\|$, and we write (with an abuse of notation) $\tilde{a}(\omega)=\tilde{a}(\|\omega\|)$ and $G(x)=G(\|x\|)$. In this case

$$
\begin{align*}
G(x) & =\frac{1}{2 \pi} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \frac{1}{\tilde{a}(\omega)} e^{i \omega \cdot x} \mathrm{~d}^{2} \omega \\
& =G(\|x\|, 0) \\
& =\frac{1}{2 \pi} \int_{0}^{\infty} \int_{0}^{2 \pi} \frac{1}{\tilde{a}(\rho)} e^{i\|x\| \rho \cos \theta} \rho \mathrm{d} \theta \mathrm{~d} \rho  \tag{11}\\
& =\int_{0}^{\infty} \frac{1}{\tilde{a}(\rho)} J_{0}(\rho\|x\|) \rho \mathrm{d} \rho
\end{align*}
$$

(the last line being an inverse Hankel or Fourier-Bessel transform [17]).
Now suppose the delta-function momentum is replaced by a rotationallyinvariant blob $f(\|x\|)$ of momentum which induces a velocity field $G_{f}$ determined by $\mathcal{A} G_{f}(x)=f(\|x\|)$. Proceeding as before gives $\widetilde{G}_{f}(\omega)=\tilde{f}(\omega) / \tilde{a}(\omega)$ and

$$
\begin{equation*}
G_{f}(x)=\int_{0}^{\infty} \frac{\tilde{f}(\rho)}{\tilde{a}(\rho)} J_{0}(\rho\|x\|) \rho \mathrm{d} \rho \tag{12}
\end{equation*}
$$

where $\tilde{f}(\omega)$ is the function of $\rho=\|\omega\|$ given by the Hankel transform of $f$ [17]:

$$
\begin{equation*}
\tilde{f}(\rho)=\int_{0}^{\infty} f(r) J_{0}(\rho r) r \mathrm{~d} r \tag{13}
\end{equation*}
$$

Comparing $\quad \widetilde{G}(\omega)=1 / \tilde{a}(\omega) \quad$ with $\quad \widetilde{G}_{f}(\omega)=\tilde{f}(\omega) / \tilde{a}(\omega)=\tilde{f}(\omega) \widetilde{G}(\omega) \quad$ shows that $G_{f}=G * f$; that is, replacing the delta function particles by blobs is the same as using delta functions with a smoothed kernel.

Finally, since the Hankel transform of a delta function is 1 , we have that considering an inertia operator with symbol $\tilde{a}(\rho)$ together with blobs of symbol $\tilde{f}(\rho)$ is identical to considering an inertia operator with $\operatorname{symbol} \tilde{a}(\rho) / \tilde{f}(\rho)$ together with delta-function particles (such operators are typically pseudo-differential). We will therefore henceforth consider, without loss of generality, delta function particles.

For numerical simulations it is convenient to have an explicit form for $G(r)$. In addition, it turns out that the dynamics of the 2-body problem depends sensitively on $G(r)$ near $r=0 . G(r)$ is negative for all $r$, reaching its global minimum at $r=0$. Thus, the velocity field induced by a single particle is in the direction $-p$. We have found it more convenient to change the sign of all kernels, so that the induced velocity field is in the direction $+p$. Since the dynamics are reversible, this does not materially affect them.

We have worked with the family of $H^{k}$ metrics with the inertia operators $\mathcal{A}_{k}$ given in equation (9) and delta-function particles. The kernels $G_{k}$ can be found explicitly in terms of modified Bessel functions of the second kind $K_{\nu}$ from equation (11) with $\tilde{a}(\rho)$ given in equation (10). The first few are (in terms of the scaled length $\tilde{r}:=r / \alpha$ )

$$
\begin{align*}
2 \pi \alpha^{2} G_{1}(r)= & K_{0}(\tilde{r}) \\
= & \log (2 / \tilde{r})-\gamma+\mathcal{O}\left(r^{2} \log r\right),  \tag{14}\\
2 \pi \alpha^{2} G_{2}(r)= & \frac{1}{2} \tilde{r} K_{1}(\tilde{r}) \\
= & \frac{1}{2}-\frac{1}{8}(2 \log (2 / \tilde{r})-2 \gamma+1) \tilde{r}^{2} \\
& -\frac{1}{128}(4 \log (2 / \tilde{r})+5-4 \gamma) \tilde{r}^{4}+\mathcal{O}\left(r^{6}\right),  \tag{15}\\
\text { and } 2 \pi \alpha^{2} G_{3}(r)= & \frac{1}{8} \tilde{r}\left(\tilde{r} K_{0}(\tilde{r})+2 K_{1}(\tilde{r})\right) \\
= & \frac{1}{4}-\frac{1}{16} \tilde{r}^{2}+\frac{1}{256}(4 \log (2 / \tilde{r})+3-4 \gamma) \tilde{r}^{4}+\mathcal{O}\left(r^{6}\right) . \tag{16}
\end{align*}
$$

Choosing the length scale $\alpha=\varepsilon / \sqrt{k}$ we have formally,

$$
\mathcal{A}_{k} \rightarrow \mathcal{A}_{\infty}:=\exp \left(-\varepsilon^{2} \nabla^{2}\right)
$$

with Fourier symbol $\tilde{a}(m)=\exp \left(-\varepsilon^{2} m^{2}\right)$ and Green's function

$$
2 \pi G_{\infty}(r)=\frac{1}{2 \varepsilon^{2}} \exp \left(-\frac{r^{2}}{4 \varepsilon^{2}}\right)
$$

That is, the Gaussian kernel is Green's function for an $H^{\infty}$ metric, the limit of a family of $H^{k}$ metrics. The family of kernels is shown in figure 1 .


Figure 1. Kernels $G_{k}(r)$ found in section 3 for a family of $H^{k}$ metrics ( $k=1,2,3,4, \infty)$ with delta-function momentum.

## 4. 2-particle systems

In this section we derive the reduced Hamiltonian for the case where there are only two particles. This enables us to describe the phase portraits, and to consider how they depend on the behaviour of the kernel, and hence on the metric.

### 4.1. The reduced Hamiltonian

Let $N=2$, so that the Hamiltonian (equation (7)) is

$$
\begin{equation*}
H=p_{1} \cdot p_{2} G\left(\left\|q_{1}-q_{2}\right\|\right)+\frac{1}{2}\left(\left\|p_{1}\right\|^{2}+\left\|p_{2}\right\|^{2}\right) G(0) \tag{17}
\end{equation*}
$$

We first change to the canonical Jacobi-Haretu center of mass coordinates [11] in order to reduce by the translational symmetry. Let

$$
d=p_{1}+p_{2}, \quad P=\left(p_{1}-p_{2}\right) / 2, \quad c=\left(q_{1}+q_{2}\right) / 2, \quad Q=q_{1}-q_{2},
$$

or

$$
p_{1}=\frac{1}{2} d+P, \quad p_{2}=\frac{1}{2} d-P, \quad q_{1}=c+\frac{1}{2} Q, \quad q_{2}=c-\frac{1}{2} Q .
$$

Then

$$
\begin{equation*}
H=\left(\frac{1}{4}\|d\|^{2}-\|P\|^{2}\right) G(Q)+\left(\frac{1}{4}\|d\|^{2}+\|P\|^{2}\right) G(0) \tag{18}
\end{equation*}
$$

Changing to symplectic polar coordinates with

$$
Q=(r \cos \theta, r \sin \theta), \quad P=\left(p \cos \theta-p_{\theta} \sin \theta / r, p \sin \theta+p_{\theta} \cos \theta / r\right)
$$

where $(r, p)$ and $\left(\theta, p_{\theta}\right)$ are conjugate variables, the reduced Hamiltonian becomes

$$
\begin{equation*}
H=\frac{1}{4}\|d\|^{2}(G(0)+G(r))+\left(p^{2}+\frac{p_{\theta}^{2}}{r^{2}}\right)(G(0)-G(r)) \tag{19}
\end{equation*}
$$

The conserved quantities are the momentum of the center of mass, $d$, and the angular momentum $p_{\theta}=: \mu$. The phase portrait of this Hamiltonian in the reduced ( $r, p$ ) phase space depends very sensitively on the behavior of $G$ near $r=0$.

First, consider the case when $G$ is smooth at $r=0$, so that $G(r)=G(0)+$ $(1 / 2) G^{\prime \prime}(0) r^{2}+(1 / 24) G^{(4)}(0) r^{4}+\mathcal{O}\left(r^{6}\right)$. In this case, near $r=0$, the Hamiltonian is

$$
\begin{equation*}
H=\frac{1}{2}\left(G(0)\|d\|^{2}-G^{\prime \prime}(0) \mu^{2}\right)+\frac{1}{24}\left(3 G^{\prime \prime}(0)\left(\|d\|^{2}-4 p^{2}\right)-G^{(4)}(0) \mu^{2}\right) r^{2}+\mathcal{O}\left(r^{4}\right) \tag{20}
\end{equation*}
$$

The $p$ axis is invariant and consists entirely of fixed points. Typically, two other orbits intersect the $p$ axis where the coefficient of $r^{2}$ in equation (20) vanishes, i.e., at

$$
\begin{equation*}
p^{2}=\frac{1}{4}\|d\|^{2}-\frac{G^{(4)}(0)}{12 G^{\prime \prime}(0)} \mu^{2} \tag{21}
\end{equation*}
$$

(Typically, and in all the examples of this article, $G^{\prime \prime}(0)<0$ and $G^{(4)}(0)>0$.) These orbits form separatrices that divide the phase space into different behavior regimes. This can be clearly seen in figure 2, where the phase portrait is given for the 2-body problem under the $\mathcal{A}_{\infty}$ Gaussian kernel. The phase portraits of the 2-body problem under other metrics are given in figures 4 and 7 and discussed in section 4.2. The energy on the $p$-axis, and hence on the separatrices, is $(1 / 2)\left(G(0)\|d\|^{2}-G^{\prime \prime}(0) \mu^{2}\right)$. As $r \rightarrow \infty, G(r) \rightarrow 0$ so $H \sim G(0)\left(p^{2}+(1 / 4)\|d\|^{2}\right)$ and the separatrices have horizontal asymptotes at

$$
\begin{equation*}
p^{2}=\frac{1}{4}\|d\|^{2}-\frac{G^{\prime \prime}(0)}{2 G(0)} \mu^{2} \tag{22}
\end{equation*}
$$

There are three principal behaviours of pairs of particles:
Scattering Both particles are free as $t \rightarrow \pm \infty$
Capture Distance between particles $r \rightarrow 0$ as $t \rightarrow \infty$
Ejection Distance between particles $r \rightarrow 0$ as $t \rightarrow-\infty$


Figure 2. Phase portrait for the 2-body problem with Gaussian kernel $G(r)=$ $\exp \left(-r^{2} / 4\right)$. The Hamiltonian is given by equation (19). Top: nonzero linear momentum $d$. Bottom: zero linear momentum. Left: nonzero angular momentum $\mu$. Right: zero angular momentum, for which the motion is confined to a line.
as well as degenerate capture/ejection orbits given by the separatrices. Inside the separatrices the two particles scatter, while outside the separatrices they either capture $\left(p G^{\prime \prime}(0)<0\right)$ or eject $\left(p G^{\prime \prime}(0)>0\right)$ each other.

We now consider the criterion for a capture to take place. Equation (22) tells us that two particles initially very far apart will capture each other if $p^{2}>(1 / 4)\|d\|^{2}-$ $\left(G^{\prime \prime}(0)\right) /(2 G(0)) \mu^{2}$. This criterion, depending on the three parameters $p,\|d\|$, and $\mu$, can be expressed more intuitively as follows. We introduce the angle of approach $\psi:=\angle p_{1} p_{2}$, the speed ratio $\left\|p_{1}\right\| /\left\|p_{2}\right\|$, and the separation distance $D . D$ is the distance of closest approach of two particles initially at $q_{1,2}$ moving in straight lines with velocities $p_{1,2}$. Clearly, as $D$ increases then the particles cease to interact, so we will seek the critical value of $D$ for a capture to occur.

We have $D=\|Q\| \sin (\angle Q P)$ and $\mu=Q \times P=\|Q\|\|P\| \sin (\angle Q P)=D\|P\|$. As $r \rightarrow \infty, p^{2} \rightarrow\|P\|^{2}=(1 / 4)\left\|p_{1}-p_{2}\right\|^{2}$, so the capture criterion becomes

$$
\begin{equation*}
-p_{1} \cdot p_{2}=-\left\|p_{1}\right\|\left\|p_{2}\right\| \cos \psi>-\frac{G^{\prime \prime}(0)}{2 G(0)} D^{2}\|P\|^{2} \tag{23}
\end{equation*}
$$

or, using the law of cosines $\|2 P\|^{2}=\left\|p_{1}-p_{2}\right\|^{2}=\left\|p_{1}\right\|^{2}+\left\|p_{2}\right\|^{2}-2\left\|p_{1}\right\|\left\|p_{2}\right\| \cos \psi$,

$$
\cos \psi<\frac{\widetilde{D}}{1+2 \widetilde{D}}\left(\frac{\left\|p_{1}\right\|}{\left\|p_{2}\right\|}+\frac{\left\|p_{2}\right\|}{\left\|p_{1}\right\|}\right), \quad-\frac{1}{4}<\widetilde{D}<0
$$

where

$$
\widetilde{D}=\frac{G^{\prime \prime}(0)}{8 G(0)} D^{2} .
$$

That is, a separation distance $D$ of less than $\sqrt{-2 G(0) / G^{\prime \prime}(0)}$ is necessary for a capture to be possible. Particles with $D=0$ (equivalently, $\mu=0$ ) are captured when $|\psi|>\pi / 2$ (a 'head-on' approach) and scattered when $|\psi|<\pi / 2$ (a 'glancing' approach), for all $\left\|p_{1}\right\| /\left\|p_{2}\right\|$; increasing $D$ and $\left\|p_{1} \mid /\right\| p_{2} \|$ moving away from 1 both restrict captures to more nearly head-on approaches. Figure 3 shows the minimum required approach angle $\psi$ for each $D$.

During a capture the particles get infinitely close, $r \rightarrow 0$. Although $p \rightarrow-\infty$, the actual speed of approach of the particles, $\dot{r}$, tends to 0 . The particles do not collide in a finite time but rather become bound together, or stick to each other. For large $p$


Figure 3. Minimum required approach angle $\psi \in((\pi / 2), \pi)$ for a capture to take place for a pair of initially distant particles with speeds $\left\|p_{1,2}\right\|$ and separation distance $D$. Here we have taken $G(r)=e^{-r^{2} / 4}$, for which the maximum possible separation distance in a capture is 2 . That this equals the range of $G$ is something of a coincidence, because the critical $D$ depends only on $G(0)$ and $G^{\prime \prime}(0)$.
and small $r$, we can use the approximate form of the Hamiltonian for $r \approx 0$ to get the approximate equations of motion:

$$
\dot{r}=-G^{\prime \prime}(0) p r^{2}, \quad \dot{p}=G^{\prime \prime}(0) p^{2} r,
$$

with solution

$$
\begin{equation*}
r(t)=r_{0} e^{-\alpha t}, \quad p(t)=p_{0} e^{\alpha t} \tag{24}
\end{equation*}
$$

where the constant $\alpha=G^{\prime \prime}(0) p_{0} r_{0}$, corresponding to a capturing orbit for $p_{0} G^{\prime \prime}(0)<0$ and an ejecting orbit for $p_{0} G^{\prime \prime}(0)>0$. We call a pair of particles in the limiting state of (24) a dipole, and explore them further in section 5.

For a scattering orbit, using

$$
\dot{\theta}=\frac{\partial H}{\partial p_{\theta}}=2 \mu \frac{G(0)-G(r)}{r^{2}},
$$

we see that the scattering angle $\left.\theta\right|_{-\infty} ^{\infty}$ ranges over $(0, \infty)$ as the minimum separation of the particles decreases from $\infty$ to 0 . (As $r \rightarrow 0, \dot{\theta}=\mathcal{O}(1)$ and $\dot{p}=\mathcal{O}(r)$, so the phase $\theta$ accumulates without bound.) Again, because close particles move very slowly, there can be an arbitrarily long delay during the close approach - the particles appear to stick together for a while.

### 4.2. Particle dynamics under different metrics

We can consider different kernels (i.e., different functions $G(r)$ ) and use them to discuss the particle dynamics for each metric in turn. In all cases, we consider the Helmholtz-style inertia operator described previously: $\mathcal{A}_{k}:=\left(1-\alpha^{2} \nabla^{2}\right)^{k}$. We begin with the $H^{1}$ metric, and progress through the different $H^{k}$ metrics.

Delta-function particles for the inertia operator $\mathcal{A}_{1}\left(H^{1}\right.$ metric) are not well-defined because they do not have a finite self-induced velocity (the same situation occurs in the 3D Euler fluid equations when the vorticity is concentrated on a curved filament). To use the delta-function particles, it is necessary to regularize the equations. One approach to this would be to simply set $G(0)=0$, so that the self-induced velocity of each particle is 0 . We consider this subsequently (section 4.3), but we first consider the standard regularization of introducing a smooth blob function $f(\|x\|)$, so that we consider the operator with symbol $\tilde{a}(\rho) / \tilde{f}(\rho)$, where $\tilde{a}(\rho)$ is the Fourier symbol of the inertia operator $\mathcal{A}$, as we described in section 3 .

We consider the Gaussian blob $f(x)=(1 / \pi) e^{-x^{2} / \varepsilon^{2}} / \varepsilon^{2}$, which tends to $\delta(x)$ as $\varepsilon \rightarrow 0$. In this case, $\tilde{f}(\rho)=(1 / 2 \pi) e^{-\varepsilon^{2} \rho^{2} / 4}$. The kernel for this blob is smooth with the inertia operator $\mathcal{A}_{1}=1-\nabla^{2}$, so we apply the results of section 4.1 and expand for small $x$ as:

$$
\begin{equation*}
2 \pi G(r)=C+\left(-\frac{1}{2 \varepsilon^{2}}+\frac{1}{4} C\right) x^{2}+\left(\frac{1}{8 \varepsilon^{4}}-\frac{1}{32 \varepsilon^{2}}+\frac{1}{64} C\right) x^{4}+\mathcal{O}\left(x^{6}\right) \tag{25}
\end{equation*}
$$

where $C=(1 / 2) \exp \left((1 / 4) \varepsilon^{2}\right) \mathrm{Ei}_{1}\left((1 / 4) \varepsilon^{2}\right)$.

As $\varepsilon \rightarrow 0$ we have $G(0) \sim(-2 \log \varepsilon-\gamma+\log 4) /(4 \pi), G^{\prime \prime}(0) \sim-1 /\left(2 \pi \varepsilon^{2}\right)$, and $G^{\prime \prime \prime \prime}(0) \sim 3 /\left(2 \pi \varepsilon^{4}\right)$, and so the separatrix intersects $r=0$ at $p=\mathcal{O}(\mu / \varepsilon)$, with a horizontal asymptote at $p=\mathcal{O}\left(\mu /\left(\varepsilon(\log |\varepsilon|)^{1 / 2}\right)\right.$. As $\varepsilon \rightarrow 0$, providing that $\mu \neq 0$, the separatrix moves outward and particle captures become rarer.

The $H^{2}$ metric is smooth enough for delta-function particles to be well-defined, with a finite self-induced velocity. The phase portrait for this case is shown in figure 4. However, $G_{2}(r)$ is not twice differentiable at $r=0$ and so the above analysis of the phase portrait of the two body problem is not valid. In fact, the singularity of $G_{2}(r)$ is sufficient to prevent the occurrence of capture orbits; when $\mu \neq 0$ all orbits are scattering (see figure 4). However, two particles with large initial relative momentum $p$ can approach arbitrarily closely, and achieve arbitrarily high $p$ during the close encounter (which makes them behave a lot like capturing orbits); but they do eventually separate. A typical such encounter is shown in figure 5.

The $H^{3}$ metric is smoother, but $G_{3}(r)$ is still not four times differentiable at $r=0$, so the above separatrix analysis is still not valid. However, one can check that the


Figure 4. Phase portrait for the 2-body problem with inertia operator $\mathcal{A}_{2}=$ $\left(1-\nabla^{2}\right)^{2}$. The Hamiltonian is given in equation (19). Unless the angular momentum $\mu$ is zero (in which case the motion is confined to a line), all orbits are scattering.



Figure 5. A typical scattering orbit for the 2-body problem with inertia operator $\mathcal{A}_{2}$ (corresponding to the phase portrait in the top left plot of figure 4). The figure below shows a magnification of the point of closest interaction between the particles. The relative momentum $2 P=p_{1}-p_{2}$ can become arbitrarily large during close approaches.

2-body problem does now have a separatrix that divides capture and scattering orbits, although it is not a smooth curve at $r=0$.

For $H^{k}(k \geq 4)$ and $H^{\infty}$, the separatrix is as described earlier. Figure 2 shows the phase portrait for the $H^{\infty}$ metric. A typical scattering orbit is shown in figure 6; in contrast to the $H^{2}$ case, the relative momentum $2 P=p_{1}-p_{2}$ is bounded over all scattering orbits. In no case are there any periodic orbits in the two body problem.


Figure 6. A typical scattering orbit for the 2-body problem with inertia operator $\mathcal{A}_{\infty}$. The relative momentum $2 P=p_{1}-p_{2}$ remains bounded during the close approach.

### 4.3. An alternative regularization for $\mathbf{H}^{1}$

As was mentioned earlier, there is another possible regularization of the $N$-particle problem with the $H^{1}$ metric, which we consider briefly here. This is simply to set the self-induced velocity of each particle to zero, i.e., $G(0)=0$. This seems drastic, when the 'correct' self-induced velocity of a delta-function is infinite, but it still corresponds to a consistent discretization of the PDE (4) in the limit of a large number of particles spaced over a curve or an area. (The contribution to the velocity at a point from nearby momentum $p$ is $\int_{0}^{\varepsilon} p(r) G_{1}(r) r \mathrm{~d} r$ for a curve, and $\int_{0}^{\varepsilon} \int_{0}^{2 \pi} p(r, \theta) G_{1}(r) r^{2} \mathrm{~d} r$ for an area; in both cases the logarithmic singularity in $G_{1}(r)$ is weak enough that both integrals tend to zero as $\varepsilon \rightarrow 0$ and the nearby momentum does not contribute to the local velocity.)

However, setting $G(0)=0$ changes the character of the 2-body problem completely. It now features periodic orbits and scattering orbits, but no capturing orbits (figure 7). We initially thought that the lack of capturing orbits was a good thing and that this regularization might be suitable for simulations with large $N$ and for approximating smooth momentum distributions (as in the point vortex method). However, numerical simulations of the 3-body problem in this case (not shown)


Figure 7. Phase portrait for the 2-body problem with inertia operator $\mathcal{A}_{1}$ and regularization $G(0)=0$. The dynamics is somewhat reminiscent of traditional 2-body problems like the Kepler problem of Newtonian gravity: there are no capture orbits; for nonzero angular momentum (left) there are scattering and periodic orbits, while for zero angular momentum (right) genuine collisions are possible.
indicates that orbits can reach a singularity in a finite time, via a curious mechanism in which two particles orbit closer and closer, under regulation from the approach of the third particle. It appears that there is no long-time existence for solutions of the $N$-body problem under this regularization. Because of this singularity we have not considered the dynamics of this $N$-particle system further.

## 5. Capture orbits and dipoles

Consider the $N$-body problem together with a smooth kernel so that 2-body captures are possible. Suppose that particles 1 and 2 become close enough to capture each other and that there are no other nearby particles. We will call the limiting state of such a captured pair of particles a 'dipole'. In center-of-mass coordinates for particles 1 and 2 only, the dipole is described by its center of mass $c$, its momentum $d$, and its orientation $Q$ (equivalently $(r, \theta)$ ) with conjugate
momentum $P$ (equivalently $\left(p, p_{\theta}\right)$ ). To obtain the Hamiltonian $H_{1,2}$ of a lone dipole we take the limit of the 2-particle Hamitonian, equation (19), when $r \rightarrow 0, p \rightarrow \infty$ with $r p$ held fixed. From equation (20), this is

$$
\begin{equation*}
H_{1,2}:=-\frac{1}{2} G^{\prime \prime}(0)\left(r^{2} p^{2}+p_{\theta}^{2}\right)+\frac{1}{2} G(0)\|d\|^{2} . \tag{26}
\end{equation*}
$$

The motion of a lone dipole is given by

$$
\begin{equation*}
c(t)=c_{0}+t G(0) d_{0}, \quad d(t)=d_{0}, \quad r(t)=r_{0} e^{-\alpha t}, \quad p(t)=p_{0} e^{\alpha t}, \quad \theta(t)=\theta_{0}+\omega t \tag{27}
\end{equation*}
$$

where the frequency $\omega=-G^{\prime \prime}(0) \mu$ and $\alpha=G^{\prime \prime}(0) p_{0} r_{0}>0\left(G^{\prime \prime}(0)<0, p_{0}<0\right)$. The terms neglected in passing from the 2-particle Hamiltonian to $H_{1,2}$ are $\mathcal{O}\left(e^{-2 \alpha t}\right)$.

We will now show that the presence of other particles, i.e., $N>2$, affects all of the degrees of freedom of the dipole, causing the internal dipole parameters ( $\mu$ and $r_{0} p_{0}$ ) to evolve.

Writing $H_{1, \ldots, N}$ for the Hamiltonian of particles 1 through $N$, the terms in $H_{1, \ldots, N}$ coupling the dipole to particle $j$ are:

$$
\begin{align*}
H_{j}^{d} & :=p_{j} \cdot\left[\left(\frac{1}{2} d+P\right) G\left(\left\|c-q_{j}+\frac{1}{2} Q\right\|\right)+\left(\frac{1}{2} d-P\right) G\left(\left\|c-q_{j}-\frac{1}{2} Q\right\|\right)\right] \\
& =\left(d \cdot p_{j}\right) G\left(\left\|c-q_{j}\right\|\right)+\left(p_{j} \cdot P\right)\left(\left(c-q_{j}\right) \cdot Q\right) \frac{G^{\prime}\left(\left\|c-q_{j}\right\|\right)}{\left\|c-q_{j}\right\|} \tag{28}
\end{align*}
$$

plus terms of order $\mathcal{O}\left(e^{-\alpha t}\right)$, which we omit if the dipole is taken to be in its limiting state. The first term in $H_{j}^{d}$ is the Hamiltonian for a single particle at position $c$ with momentum $d$, while the second (dipole) term generates an $\mathcal{O}(1)$ contribution to all of $\dot{d}, \dot{p}_{j}, \dot{q}_{j}, \dot{Q} / Q$, and $\dot{P} / P$. Identifying $q_{N+1} \equiv c$ and $p_{N+1} \equiv d$, the Hamiltonian for a dipole and $N-2$ singleton particles is

$$
\begin{aligned}
H_{1, \ldots, N} & =\frac{1}{2} \sum_{i, j=1}^{N} p_{i} \cdot p_{j} G\left(\left\|q_{i}-q_{j}\right\|\right) \\
& =\frac{1}{2} H_{1,2}+\frac{1}{2} H_{3, \ldots, N+1}+\frac{1}{2} \sum_{j=3}^{N} H_{j}^{d} .
\end{aligned}
$$

This suggests that the overall evolution of $N$ particles with a smooth kernel is described by (i) free, straight-line motion of all particles when all particles are well-separated; (ii) scattering and interactions when distances are moderate; and (iii) capture or sticking together of two particles when they come within range, after which they continue to evolve as a dipole (and interact with the other particles). It is also possible for more than two particles to simultaneously capture each other. We have not analyzed this situation but it seems to follow the same course as the 2-body capture, with all inter-particle distances tending to 0 exponentially fast.

It may also be possible for the influence of the other particles to cause a dipole to separate, although we never observed this happening.

If the boundary conditions are taken to be periodic, this indicates that most orbits of the $N$-particle problem end up with all particles stuck together in one lump.

A typical orbit for 20 particles in the plane is shown in figure 8. The initial positions and velocities are (normally) randomly distributed. Here, the particles have been captured into groups of 9,3 (with a 4th about to join), 2 (with 2 more possible captures), and 3 singletons.

We now analyze the influence of a dipole on a third particle with zero momentum - a 'test particle'. The test particle does not affect the motion of the dipole. We only have to find the position $q$ of the test particle, which obeys

$$
\begin{equation*}
\dot{q}=d(t) G(\rho)+P((c(t)-q) \cdot Q) f(\rho), \tag{29}
\end{equation*}
$$

where $c(t)$ and $d(t)$ are given by equation (27), $\rho=\|c-q\|$ is the distance from the dipole to the test particle, and $f(\rho)=G^{\prime}(\rho) / \rho$.

One solution of (29) is $q(t)=c(t)$, in which the test particle sits on top of the dipole. We shall see that dipoles can capture test particles into this state.

First consider the case $d=0$ of a motionless dipole located at $c=0$. We compute

$$
\begin{aligned}
r P & =\binom{p_{0} r_{0} \cos \omega t-\mu \sin \omega t}{p_{0} r_{0} \sin \omega t+\mu \cos \omega t} \\
& =A^{-1}\binom{-\sin \beta}{\cos \beta} \\
Q & =\binom{r \cos \omega t}{r \sin \omega t}
\end{aligned}
$$

where $A^{2}=p_{0}^{2} r_{0}^{2}+\mu^{2}$ and $\beta=\omega t+\phi$, and the phase $\phi$ is defined by $\cos \phi=\mu / A$, $\sin \phi=-p_{0} r_{0} / A$. Equation (29) becomes

$$
\dot{q}=-A^{-1}\left(q_{1} \cos \omega t+q_{2} \sin \omega t\right) f(\rho)\binom{-\sin \beta}{\cos \beta}
$$

note the $e^{\alpha t}$ factors cancel out and the interaction is characterized by the parameters $r_{0} p_{0}$ and $\mu$ of the dipole. The form of these equations suggests going into a rotating frame by

$$
y=\left(\begin{array}{cc}
\cos \beta & \sin \beta \\
-\sin \beta & \cos \beta
\end{array}\right) q,
$$

which leads to a drastic simplification to the planar, autonomous system

$$
\begin{align*}
& \dot{y}_{1}=\omega y_{2}  \tag{30}\\
& \dot{y}_{2}=-\left(\mu y_{1}+r_{0} p_{0} y_{2}\right) f(\rho)-\omega y_{1} .
\end{align*}
$$

Note $\rho=\|q\|=\|y\|$. The phase portrait of this system is shown in figure 10 .





$T=1.3$



$T=2.2$










Figure 8. Typical orbit for 20 randomly placed particles under a Gaussian kernel. (Top: particle paths, with final particle positions indicated by a dot; bottom: snapshots of the particle positions at 16 different times.)

We now investigate the stability of the fixed point $y=0$ under the assumption that $G$ is smooth, so that $f(\rho)=G^{\prime \prime}(0)+(1 / 6) G^{(4)}(0) \rho^{2}+\cdots$ and (30) becomes

$$
\begin{align*}
& \dot{y}_{1}=\omega y_{2}=-G^{\prime \prime}(0) \mu y_{2}  \tag{31}\\
& \dot{y}_{2}=r_{0} p_{0} G^{\prime \prime}(0) y_{2}-\frac{1}{6} G^{(4)}(0)\left(y_{1}^{2}+y_{2}^{2}\right)+o\left(\|y\|^{2}\right) \tag{32}
\end{align*}
$$

The eigenvalues of the fixed point $y=0$ are 0 (eigenvector $(1,0)^{T}$ ) and $-p_{0} r_{0} G^{\prime \prime}(0)<0$ (eigenvector $\left(\mu, p_{0} r_{0}\right)^{T}$ ). The fixed point is nonhyperbolic, but a routine application of center manifold theory then gives a center manifold located at $y_{2}=\gamma y_{1}^{3}+\mathcal{O}\left(y_{1}^{4}\right)$ with $\gamma=-\mu G^{(4)}(0) /\left(6 p_{0} r_{0} G^{\prime \prime}(0)\right)$, and, on the center manifold, the dynamics is governed by the reduced equation

$$
\dot{y}_{1}=\delta y_{1}^{3}, \quad \delta=\omega \gamma=\frac{\mu^{2} G^{(4)}(0)}{6 p_{0} r_{0}} .
$$

Recalling that $p_{0}<0$ for a dipole, we have that $y=0$ is asymptotically stable if $G^{(4)}(0)>0$ and unstable if $G^{(4)}(0)<0$. The fixed point is stable for the Gaussian kernel $G(\rho)=e^{-\rho^{2}}$.

That is, the dipole for a Gaussian kernel attracts nearby test particles, albeit very slowly, since $y_{1}(t)=\left(y_{1}(0)^{-2}-2 \delta t\right)^{-1 / 2}$. A typical test particle far from the dipole moves initially slowly (since the velocity field is exponentially small there), then quickly falls onto the centre manifold, then orbits the dipole at the same frequency as the dipole itself, maintaining a phase shift of $\phi$, at a gradually diminishing distance (figures 9 and 10).

Finally, we consider a moving dipole, which we take (e.g., by rotating space and by scaling time) to be moving along the $x$-axis at unit speed, $c(t)=t(1,0)$. A test particle with sufficiently large $\left|q_{2}(0)\right|$ (i.e., with the test particle far enough from the path of the dipole) will not be entrained as the dipole passes. However, the critical value of $q_{2}(0)$ will also depend on $q_{1}(0)$, which determines the phase of the dipole as


Figure 9. Attraction of a test particle to a dipole at the origin. \{The particle moves [under equation (28)] in the direction of the rotating direction vector of the dipole, with the cusps corresponding to the dipole pointing directly at the particle.\}


Figure 10. Phase portrait of equation (30) for a test particle in the field of a motionless dipole in rotating coordinates. Here, $\mu=1, r_{0} p_{0}=-1$, and $G(\rho)=e^{-\rho^{2}}$.
it approaches. A numerical simulation (figure 11) reveals a surprising asymmetry where, for an anticlockwise $(\mu>0)$ dipole, the entrainment is independent of the phase for $q_{2}(0)<0$ and strongly dependent on the phase for $q_{2}(0)>0$. All particles out to $q_{2}=-1.536$, on the right of the dipole, are entrained, but on the left the critical initial distance varies from 0.6 to 1.15 .

## 6. Discussion

The particle dynamics in $\operatorname{Diff}\left(\mathbb{R}^{2}\right)$ with a $H^{k}$ metric $(k \geq 2)$ that we have considered in this article show striking contrasts to the more familiar dynamics of point vortices for $\operatorname{Diff}_{\text {vol }}\left(\mathbb{R}^{2}\right)$ with an $L^{2}$ metric. Of particular note are:

- our particles are vector particles, point vortices are point particles;
- for our particles, the 2-body problem is integrable, for point vortices the 3-body problem is integrable;


Figure 11. Entrainment of a field of test particles by a moving dipole. The dipole (whose direction is indicated by a stick) is moving to the right with speed 1. Particles shown as small dots are entrained by the dipole, while particles shown as larger dots are left behind.

- our particles exhibit short-range interactions $u(r) \sim e^{-r} r^{k-1 / 2}(r \rightarrow \infty)$, point vortices exhibit long-range interactions $u(r) \sim 1 / r$;
- our particles have no periodic 2-body orbits, whereas most 2-body orbits are periodic for point vortices; and
- the 2-body dynamics of $\operatorname{Diff}\left(\mathbb{R}^{2}\right)$ depend sensitively on the metric, whereas for point vortices, the qualitative 2-body dynamics are independent of the metric.

Clearly, we have only scratched the surface of the rich dynamics of the system (8). Further dynamical questions to be considered include

- a full study of the 3-body problem;
- determination of the long-time existence of solutions;
- a classification of the limiting states of the $N$-body problem;
- a study of the limiting (' $N$ pole') state of the capture of $N$ particles; and
- a consideration of the geometry of the Riemannian manifold whose geodesics are governed by (8),
while broader questions include
- the significance of the sensitivity to the metric for image matching applications;
- the significance of particle dynamics for solutions of the $\operatorname{PDE}(1,2)$; and
- the convergence of the point particle approximation to solutions of the PDE (1,2), including the role of regularization of the metric (especially in the $H^{1}$ case).


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